

## ON THE FIELD ANOMALY OF NEAR WAKES IN A COLLISIONLESS PLASMA \*

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A variational approach is proposed to determine the induced electric field by charge separation in the near wake of a large negatively charged body moving at mesothermal speeds in a tenuous plasma. The presence of potential well is discussed.

The interaction between the ionospheric medium and a rapidly moving body provides a unique plasma phenomenon of much geophysical consequences [1]. The moving body in question, e.g., a space probe or meteorite, can be characterized by the following conditions:  $\lambda_D \ll R \ll l$ ;  $c_i \ll V \ll c_e$  where  $R$  and  $V$  denote the size and the speed respectively of the body;  $\lambda_D$ ,  $l$ ,  $c_i$  and  $c_e$ , the Debye length, mean free path, ion thermal speed and electron thermal speed respectively of the plasma. It is well known that when a neutral plasma meets a moving blunt body a potential sheath having a thickness of the order  $\lambda_D$  develops to shield the frontal body. In the near wake behind a rapidly moving body where the electrons lead the ions in filling the void region, a charge separation field in the form of a potential valley is created. In other words, the potential distribution in the near wake, instead of varying monotonically from its boundary value at the body surface  $\phi_s$  (usually  $< 0$ ) to zero at the free stream, dips into a well of depth  $|\phi_{\min}| > |\phi_s|$  before it rises to its ambient value (zero). This anomalous field configuration which has been shown to cause plasma trapping [2] and possibly other consequences yet to be explored is of interest in the present note.

Consider a conducting spherical body of radius  $R$  and surface potential  $\phi_s$  ( $< 0$ ) which moves at a steady mesothermal speed  $V$  in a bithermal neutral plasma of singly charged ions at temperature  $T_i$  and free electrons at temperature  $T_e$  where  $\lambda_D \ll R \ll l$ . It is assumed that in collision with the body an electron is absorbed; an ion, neutralized and re-emitted as a neutral particle. The magnetic field effect is neglected. These conditions hold

e.g., for a typical satellite-motion in the upper ionosphere [3]. In the following analysis the origin of a cylindrical coordinate is fixed to the center of the moving body with its  $z$ -axis aligned with the axis of the wake. The linear displacement  $r$  is normalized by  $R$ ; field potential  $\phi(r)$  by  $e/kT_i$  where  $e$  and  $k$  denote the electronic charge and Boltzmann constant respectively; particle velocity  $c$  by  $c_i$  and the particle density  $n(r)$  by  $n_\infty$ , the free stream electron or ion density. The conditions  $\phi_s < 0$  and  $V \ll c_e$  imply that the electron distribution in the wake will deviate from its free stream value only by the Boltzmann factor thus  $n_e(r) = \exp(\phi T_i/T_e)$ . On the other hand since  $V \gg c_i$  the ion distribution  $f(r, c)$  in the wake will differ strongly with its free stream equilibrium state. In fact  $f(r, c)$  and  $\phi(r)$  are governed by the following coupled equations of the collision free (particle) equation and the Poisson (field) equation [3]:

$$c \cdot \nabla_r f - \frac{1}{2} (\nabla_r \phi) \cdot \nabla_c f = 0; (\lambda_D/R)^2 \nabla_r^2 \phi = n_i - n_e \quad (1)$$

where  $n_i = \int f dc$  with  $\phi = \phi_s$  at  $r = R$  and  $\phi = 0$  at  $r \rightarrow \infty$ .

An iterative numerical scheme has often been used [3] to treat the nonlinear system (1). It starts with a formal solution for  $f(r, c)$  from the quasi-linear first order particle equation, in terms of the appropriate invariants. Notice that an analogous mathematical formalism in stellar dynamics is known as Jeans theorem [4]. For an axi-symmetric field herein interested three isolating invariants are needed. The energy invariant  $E = c_\rho^2 + c_\theta^2 + c_z^2 + \phi$  and angular momentum invariant  $L_z = |r \times c|_z$  are obvious [3]; the existence of the third invariant for a general arbitrary axi-symmetric  $\phi(r)$ , however, has not been settled. Nonetheless, if the potential is prescribed as  $\phi_L = \zeta(r) + \eta(\theta)/r^2$  where  $\zeta(r)$  and

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$\eta(\theta)$  are arbitrary functions of the radius vector  $r$  and the polar angle  $\theta$  respectively, it is known that the third invariant  $I_L = |\mathbf{r} \times \mathbf{c}|^2 + \eta(\theta)$  [4].

It is suggested herein to use a variational method that yields approximate solutions to the system (1). It consists of finding a functional  $J(\phi, f)$  whose variation gives the equations (1) as its "Euler equations" [5]. This is found to have the following form:

$$J(\phi, f) = \frac{1}{2} (\lambda_D/R)^2 \left\{ \int (\nabla \phi)^2 d\mathbf{r} + \phi_s \int_{\text{sphere}} [\partial \phi(1, \theta)/\partial r] d\theta \right\} - \iint (n_i - n_e) d\phi' d\mathbf{r} + \iint [c \cdot \nabla_r f + (\nabla_r \phi) \cdot \nabla_c f] f' d\mathbf{c} d\mathbf{r} \quad (2)$$

for an axi-symmetric wake. Trial functions for  $\phi(\mathbf{r})$  and  $f(\mathbf{r}, \mathbf{c})$  must be furnished to functional (2) before a variational principle can be applied. It is intended to introduce a perturbed pair  $(\phi, I_3)$  which is related to the unperturbed  $(\phi_L, I_L)$ -pair as follows \* [5]:

$$\phi = \phi_L + \eta_1(r, \theta)/r^2, \quad I_3 = I_L + \eta_1(r, \theta) \quad (3)$$

$I_3$  will be used along with  $E$  and  $L_z$  to prescribe the trial  $f(\mathbf{r}, \mathbf{c}) = f(E, L_z, I_3)$  following Jeans theorem. In compliance with (3), the suggested trial potential  $\phi(\mathbf{r})$  which satisfies the wake boundary conditions is prescribed as follows:

$$\phi = \{\phi_s - (r-1) \cos^2 \theta [B_0 p_0 + r B_1 p_1 \exp -(r-1) + r^2 B_2 p_2 \exp -2(r-1) + r^i B_i p_i \exp i(r-1) + \dots]\} \exp -(r-1) \quad (4)$$

and its associated third invariant:

$$I_3 = r^2 (c_\theta^2 + c_\phi^2) - r^2 (r-1) \cos^2 \theta [B_0 + r B_1 p_1 \exp -(r-1) + r^2 B_2 p_2 \exp -2(r-1) + r^i B_i p_i \exp -i(r-1) + \dots] \exp -(r-1) \quad (5)$$

where  $p_i (i = 0, 1, 2, 3, \dots)$  are Legendre polynomials;  $B_i (i = 0, 1, 2, \dots)$ , coefficients to be determined with variational method. The trial  $f(E, L_z, I_3)$  satisfying its boundary conditions [3] can be written

$$f = \frac{1}{2} \pi^{-3/2} [1 + \text{sign}(S-1)] \exp [-(E-V^2)] \quad (6)$$

\* A priori restriction thus implied is that  $\phi$  deviates only slightly from  $\phi_L$ .

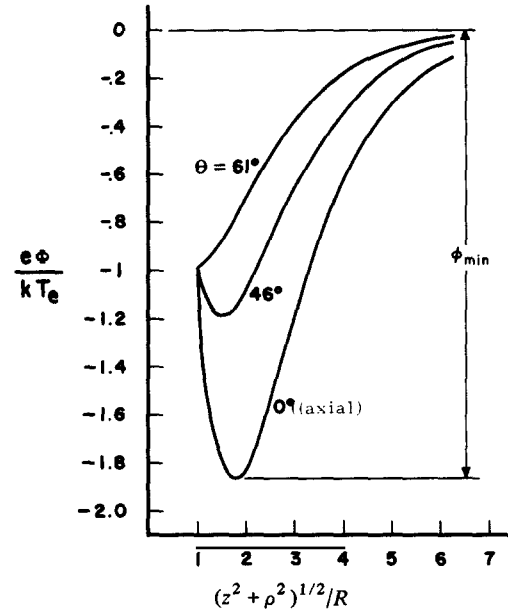


Fig. 1. Electric potential distribution in an axi-symmetric wake.

where  $S = I_3/I_3$  (evaluated at  $\theta = \pi/2$ ). The vanishing of the differential coefficients  $\partial J/\partial B_i$  in the first variant  $\delta J$ , after the substitution of eqs. (4), (5), and (6) in the functional (2), yields a set of nonlinear transcendental algebraic equations which can be solved by the generalized Newton's method. The result,  $\phi(\mathbf{r})$ , of a case with  $\phi_s = -1$ ,  $V/c_i = 8$ ,  $\lambda_D/R = 0.05$  and  $T_e = T_i$  is illustrated in fig. 1. In this illustrated variational calculation, the terms beyond  $B_3$  and  $p_3$  in eqs. (4) and (5) have been truncated. Calculations including the  $B_4$  and  $p_4$  terms have also been made with, however, negligible change in results of  $\phi(\mathbf{r})$  and  $n_i(\mathbf{r})$ . The present result (fig. 1) agrees well with previous numerical iteration results in values  $\phi_{\min}$  but shows a slight shift in  $\phi$ -distribution [3]. The locus of the potential minimum ( $\phi_{\min}$ ) forms a conical surface in an axi-symmetric wake. The presence of a potential valley in the wake stems from the charge separation as the ambient electrons and ions move into the wake with unequal mass motions.

- [1] V.C. Liu, On interplanetary gas dynamics. Adv. App. Mech. 12 (Acad. Press 1972) p. 195.
- [2] V.C. Liu, Nature 215 (1967) 127.
- [3] V.C. Liu, Space Sci. Rev. 9 (1969) 423.
- [4] D. Lynden-Bell, M.N. Royal Astron. Soc. 124 (1962) 1, 95.
- [5] M. Becker, The principles and applications of variational methods (MIT Press, 1964).